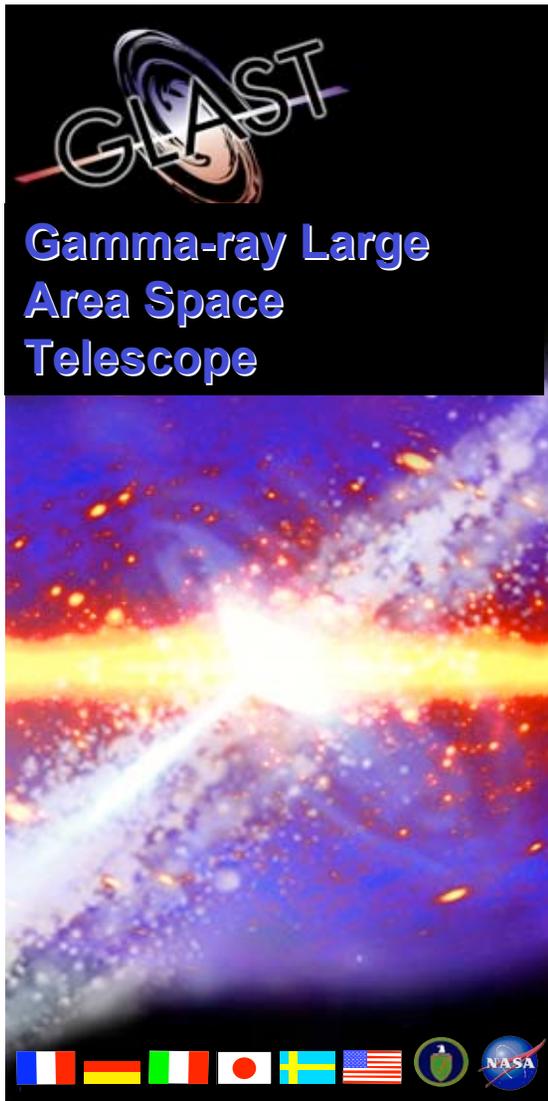




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Photon Event Maps

- Source Detection
- Transient Detection

Jeff Scargle

Space Science Division

NASA Ames Research Center

Thanks: Jay Norris, and AISRP

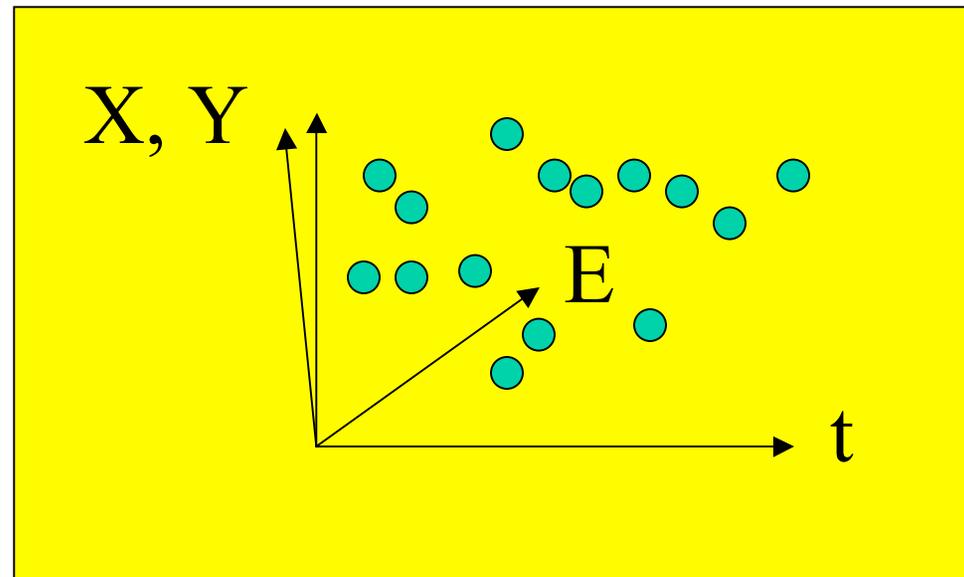


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The GLAST Data Stream

4-Dimensional Data Space:

- position on the sky
- time of arrival
- energy



$$\{ X_i, Y_i, t_i, E_i ; i = 1, 2, 3, \dots N \}$$



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Density Estimation + Structure Identification

Many analysis problems can be treated with a two-step procedure:

- Estimate photon density in the data space
- Identify and characterize structures in the density profile

Photon density estimates radiation intensity.



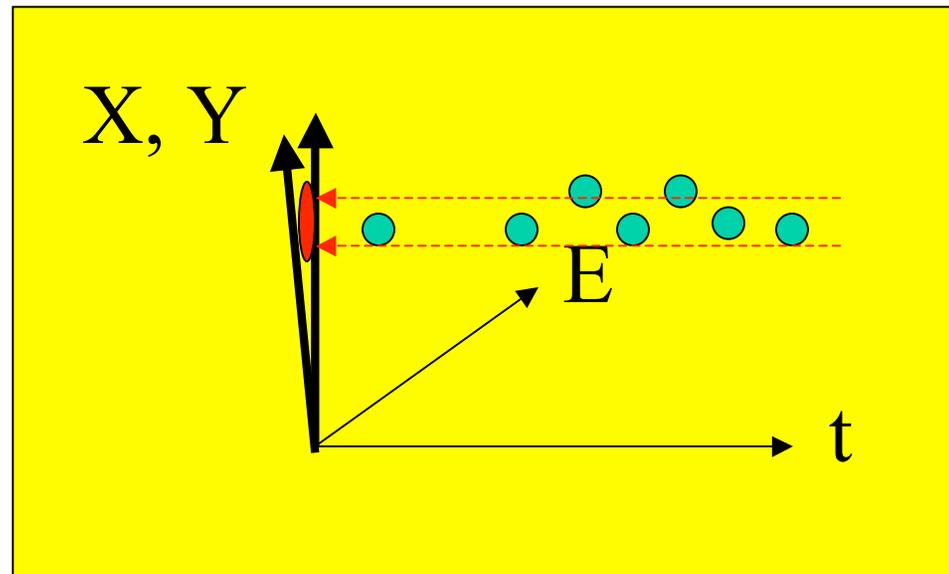
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Density Estimation + Structure Identification

Many analysis problems can be treated with a two-step procedure:

- Estimate photon density in the data space
- Identify and characterize structures in the density profile

Example:
Source
Detection
(point or
Extended)





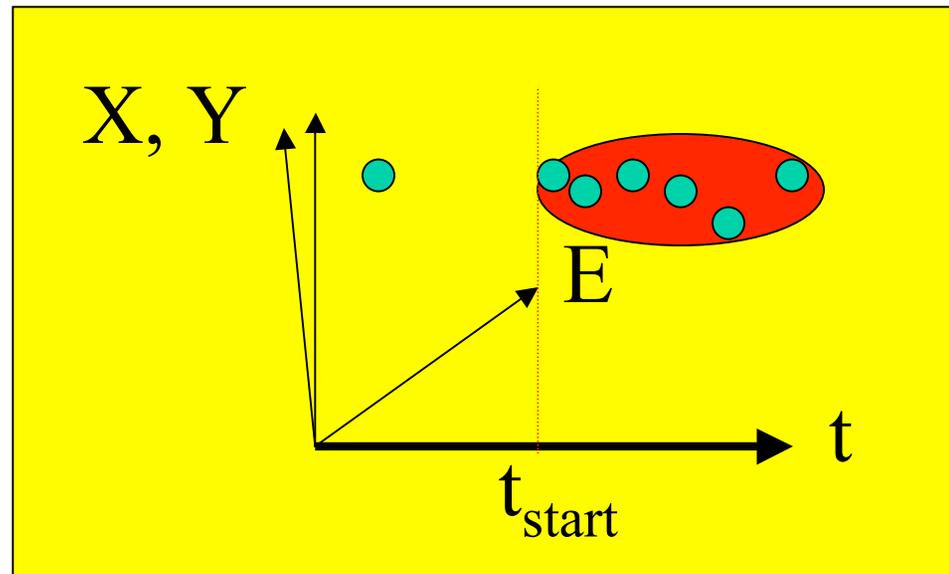
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Density Estimation + Structure Identification

Many analysis problems can be treated with a two-step procedure:

- Estimate photon density in the data space
- Identify and characterize structures in the density profile

Example:
Transient
Detection





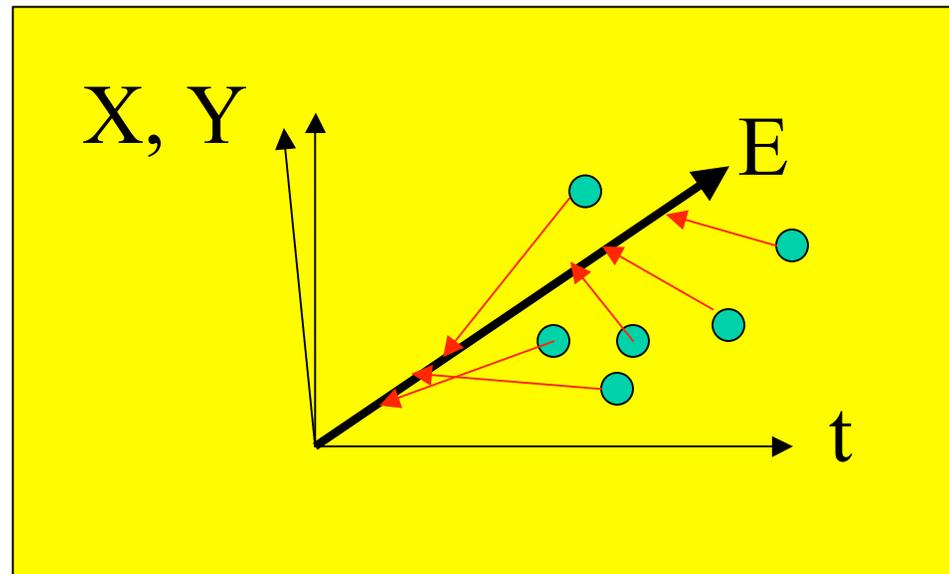
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Density Estimation + Structure Identification

Many analysis problems can be treated with a two-step procedure:

- Estimate photon density in the data space
- Identify and characterize structures in the density profile

Example:
Spectrum
Analysis





The Bin Myths

- I. Point data must be binned in order to make sense out of them.

- II. The bins must be large enough so that each bin has a significantly large sample.

The analysis described here uses no binning:

~~No spatial bins (healpix)~~

~~No spatial smoothing (such as convolution with a kernel)~~

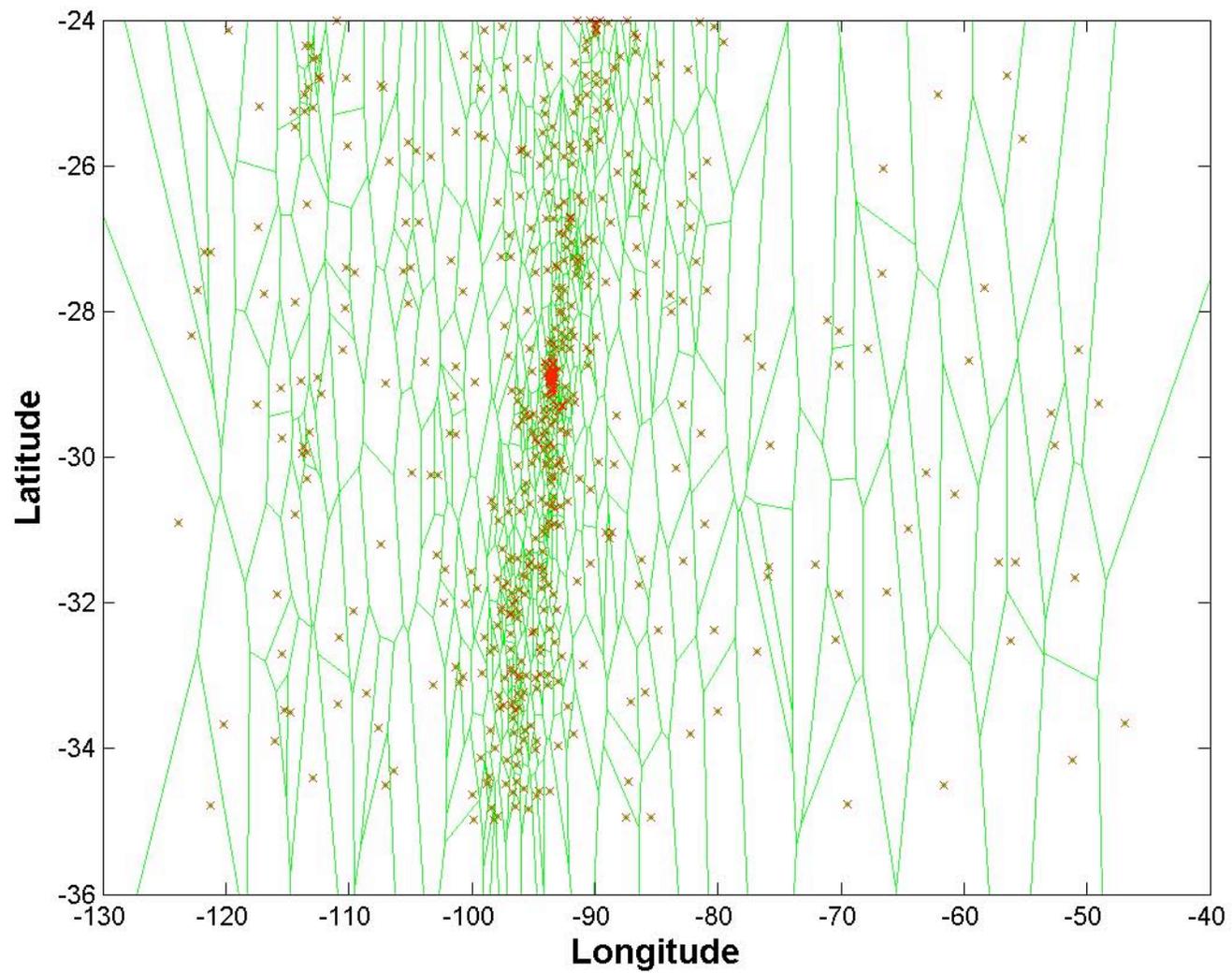
~~No sliding templates (such as likelihood test statistic)~~

~~No time bins~~

~~No energy bins~~



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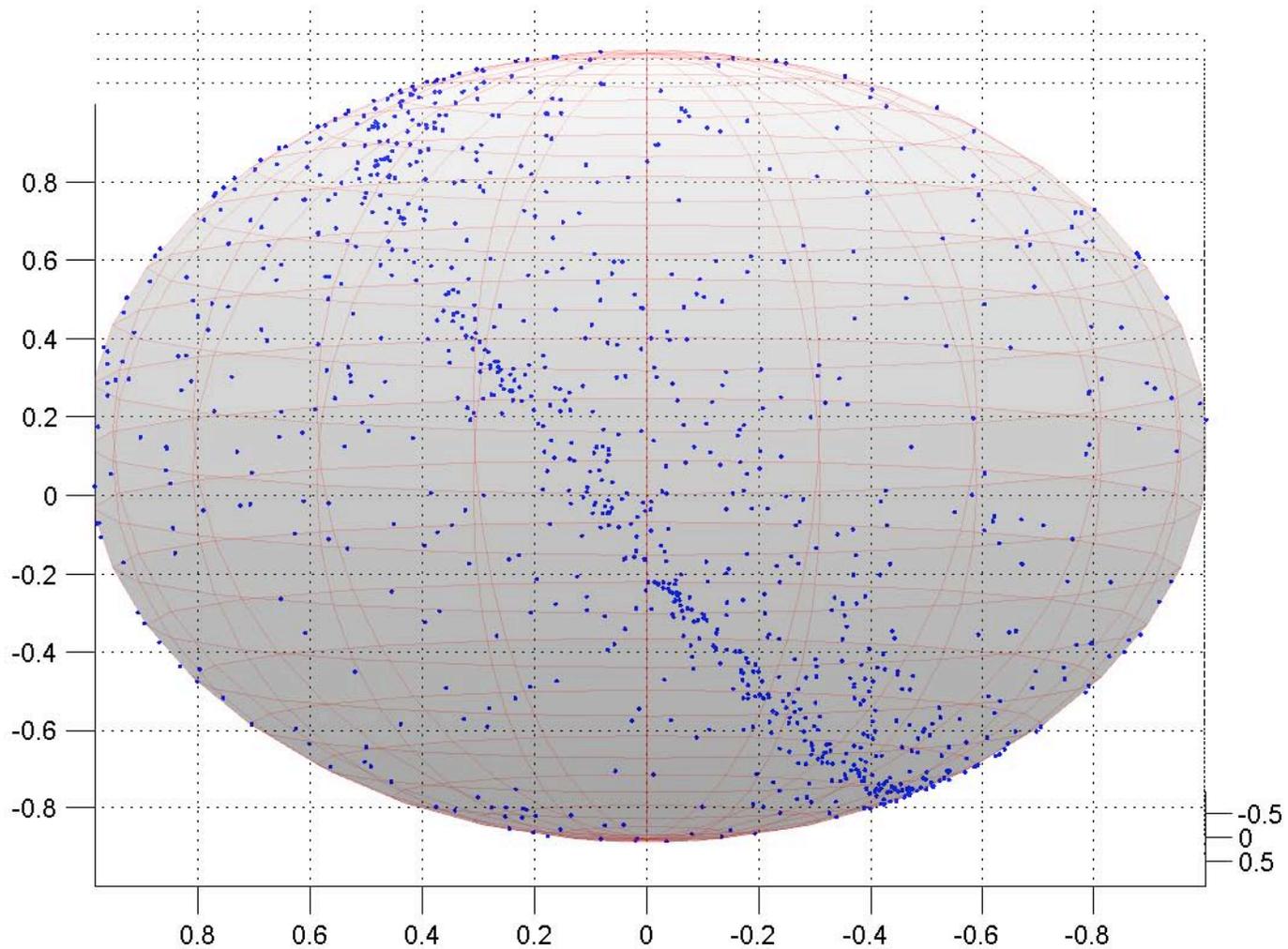


The First International GLAST Symposium – Astrostatistics Session, February 7, 2007



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Photons on the Sphere

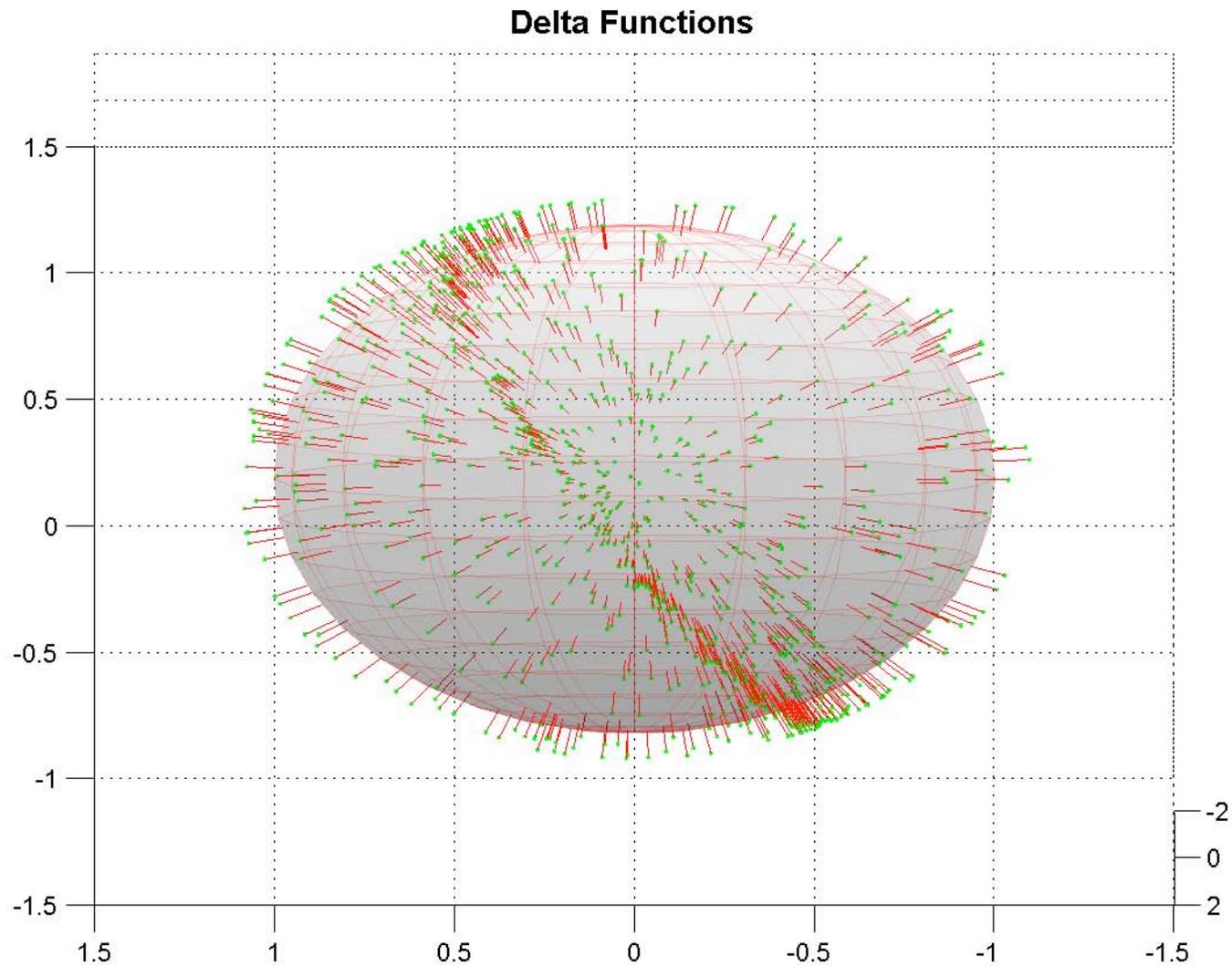


The First International GLAST Symposium – Astrostatistics Session, February 7, 2007



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Photons on the Sphere



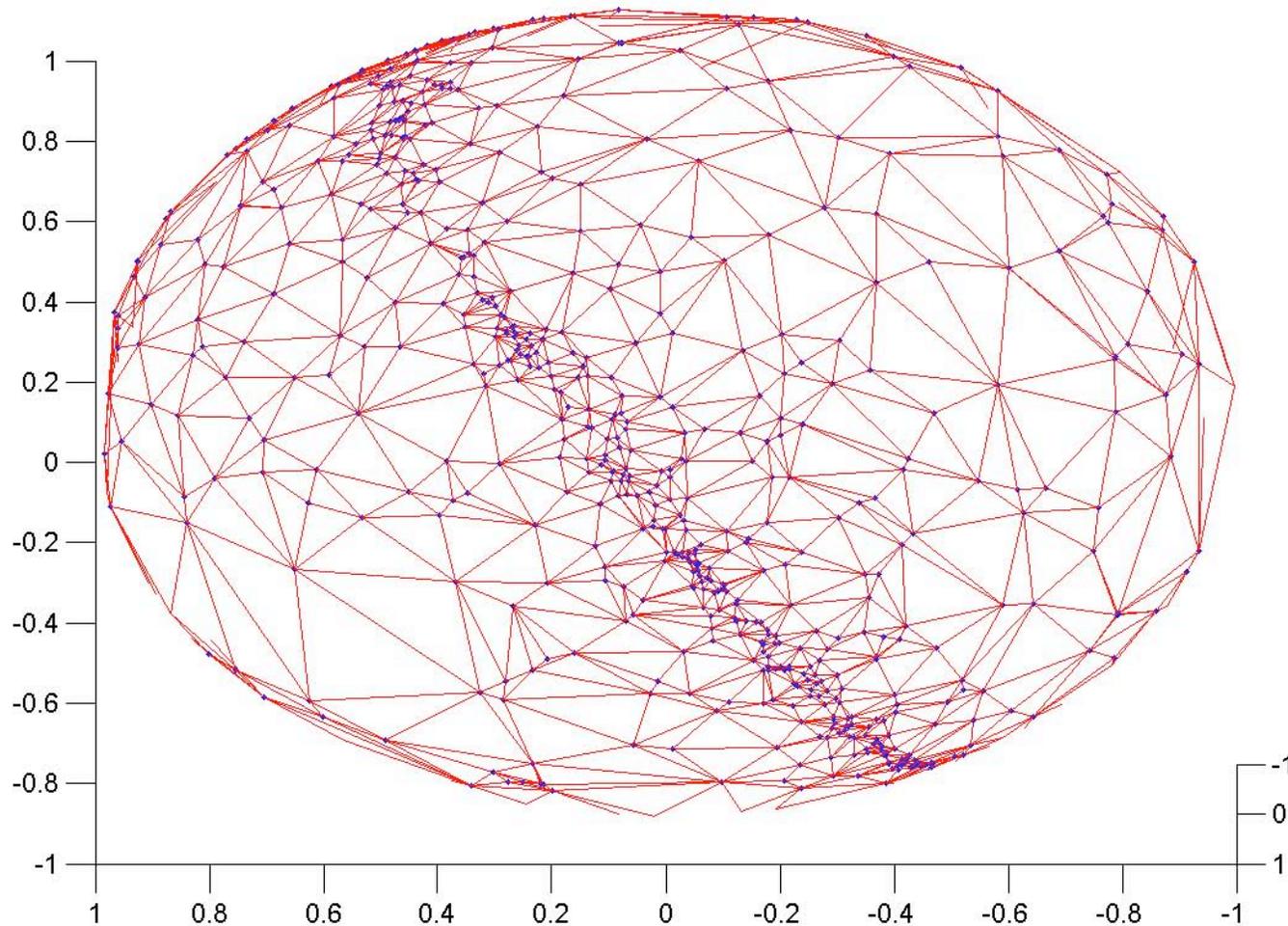
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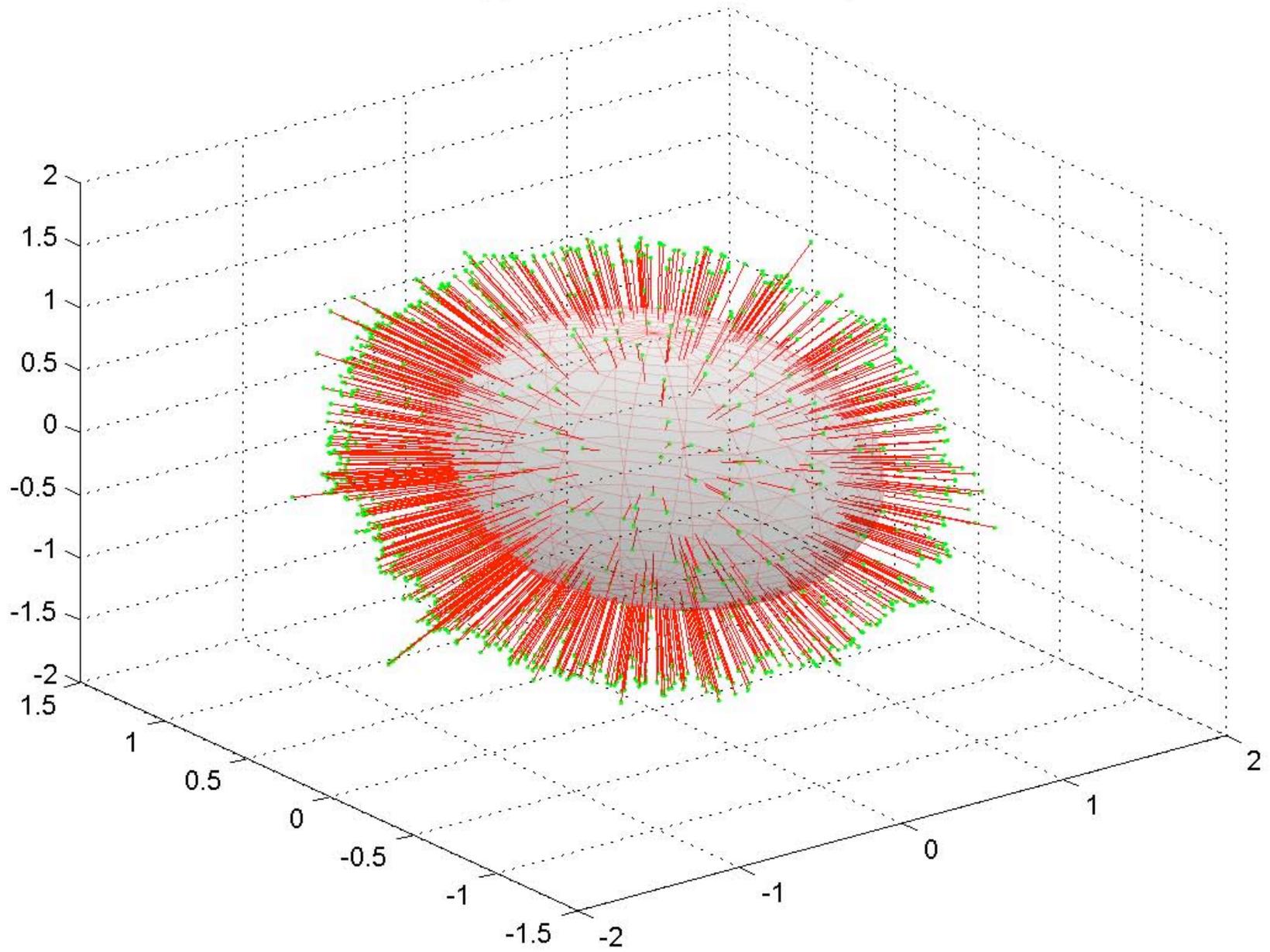
Photons on the Sphere

Photon positions on sphere \rightarrow convex hull \rightarrow Delaunay triangulation \rightarrow Voronoi tessellation



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log(1 / Voronoi Volume)

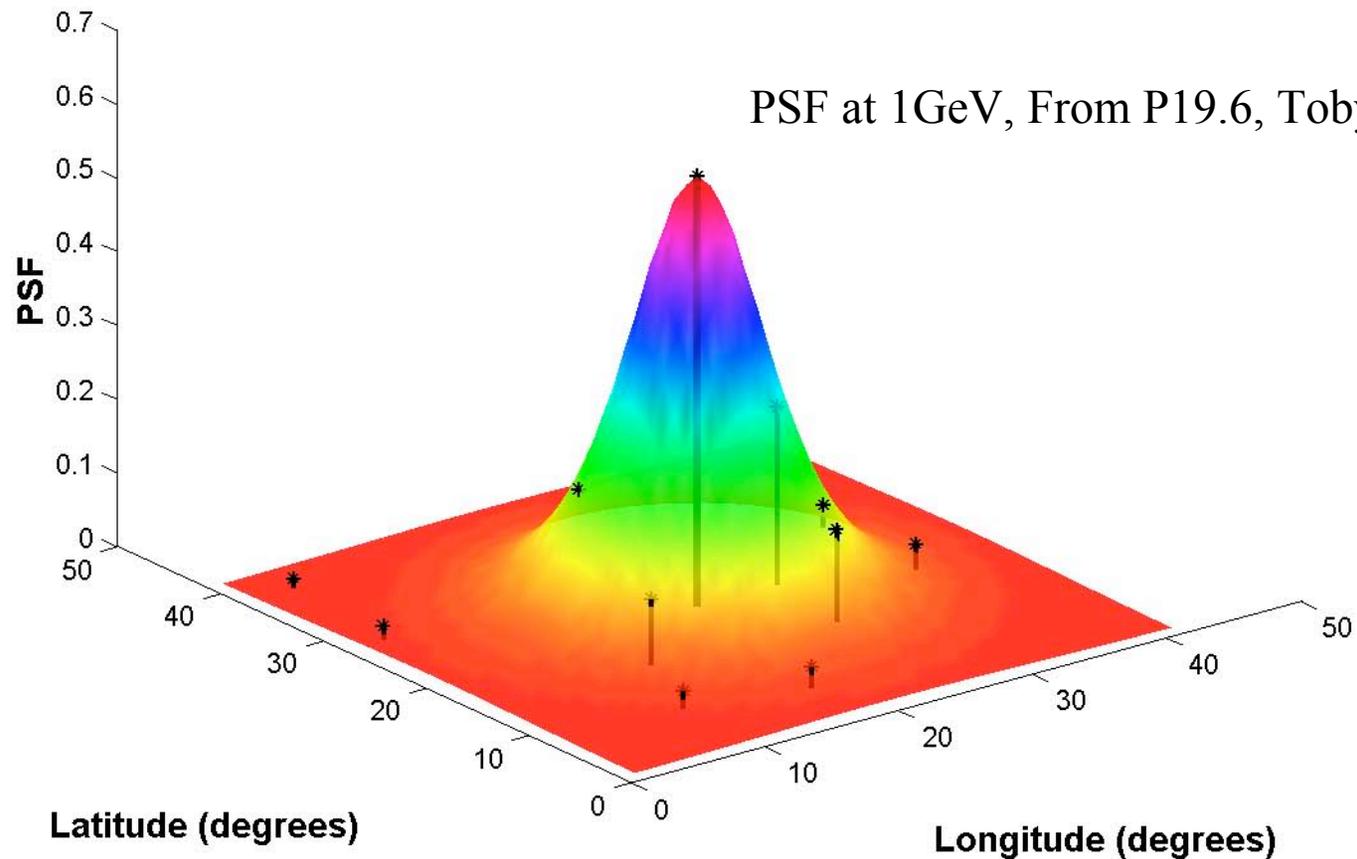




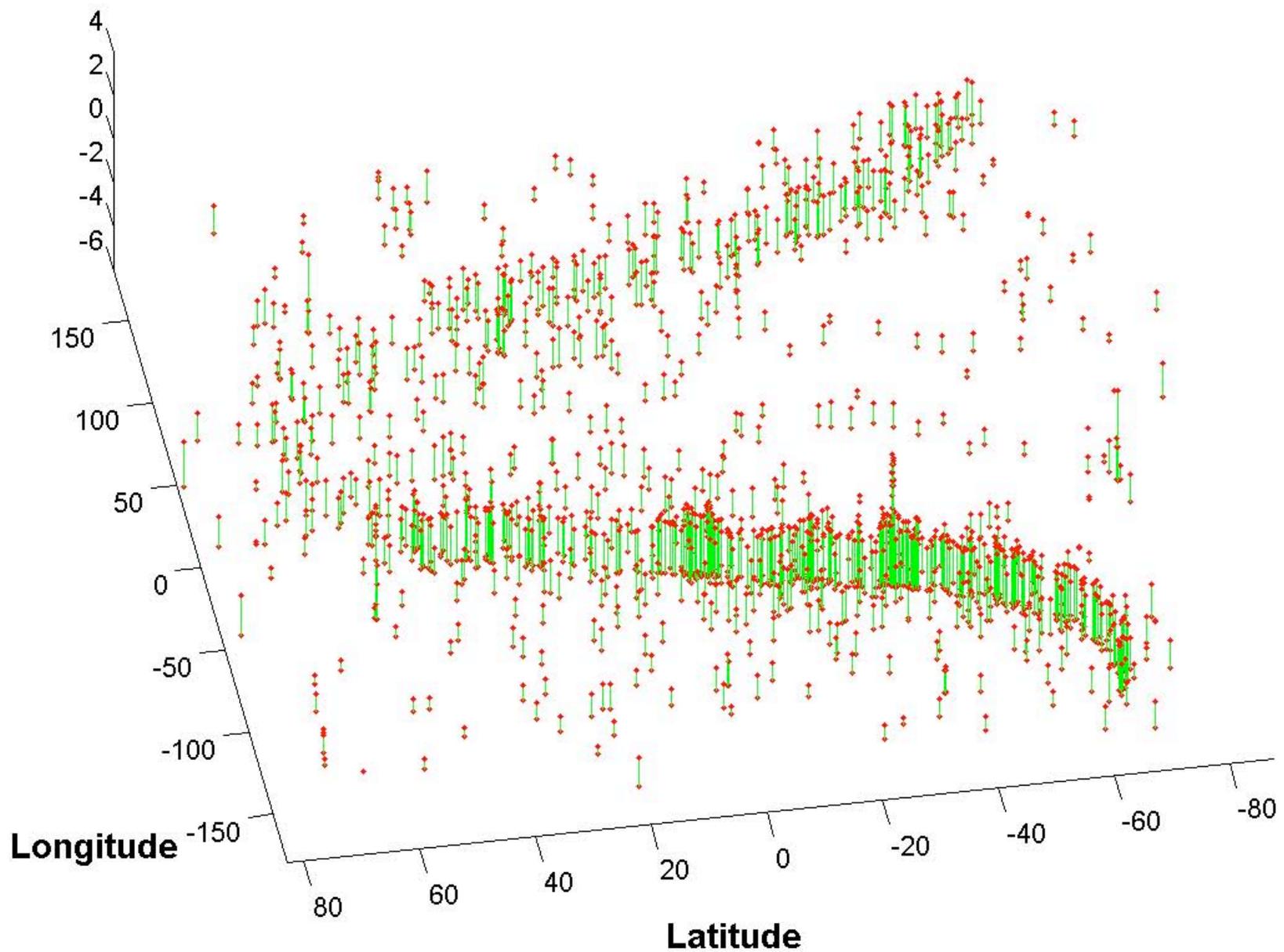
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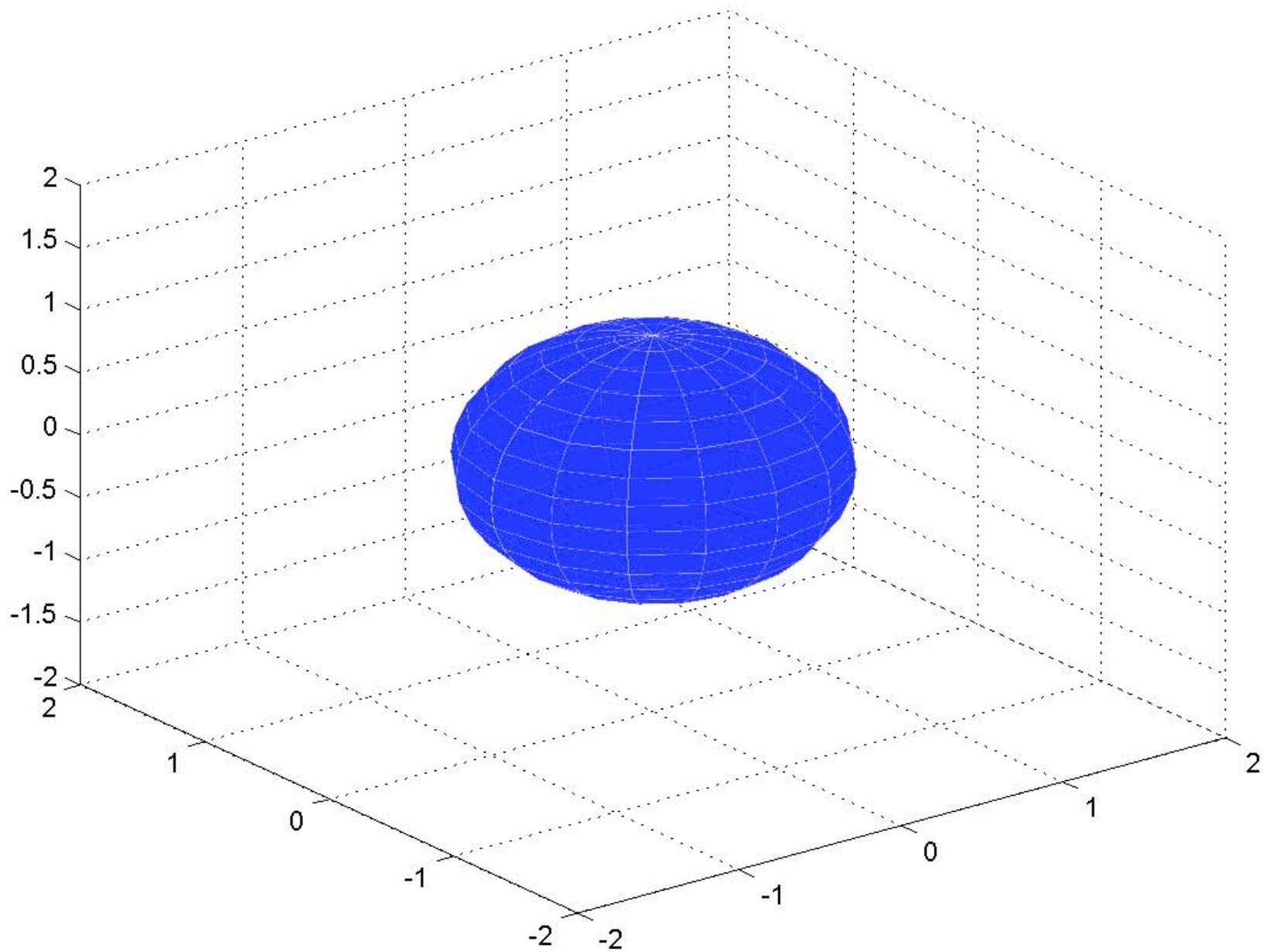
Apportion Weights to Each Nearby Photon.

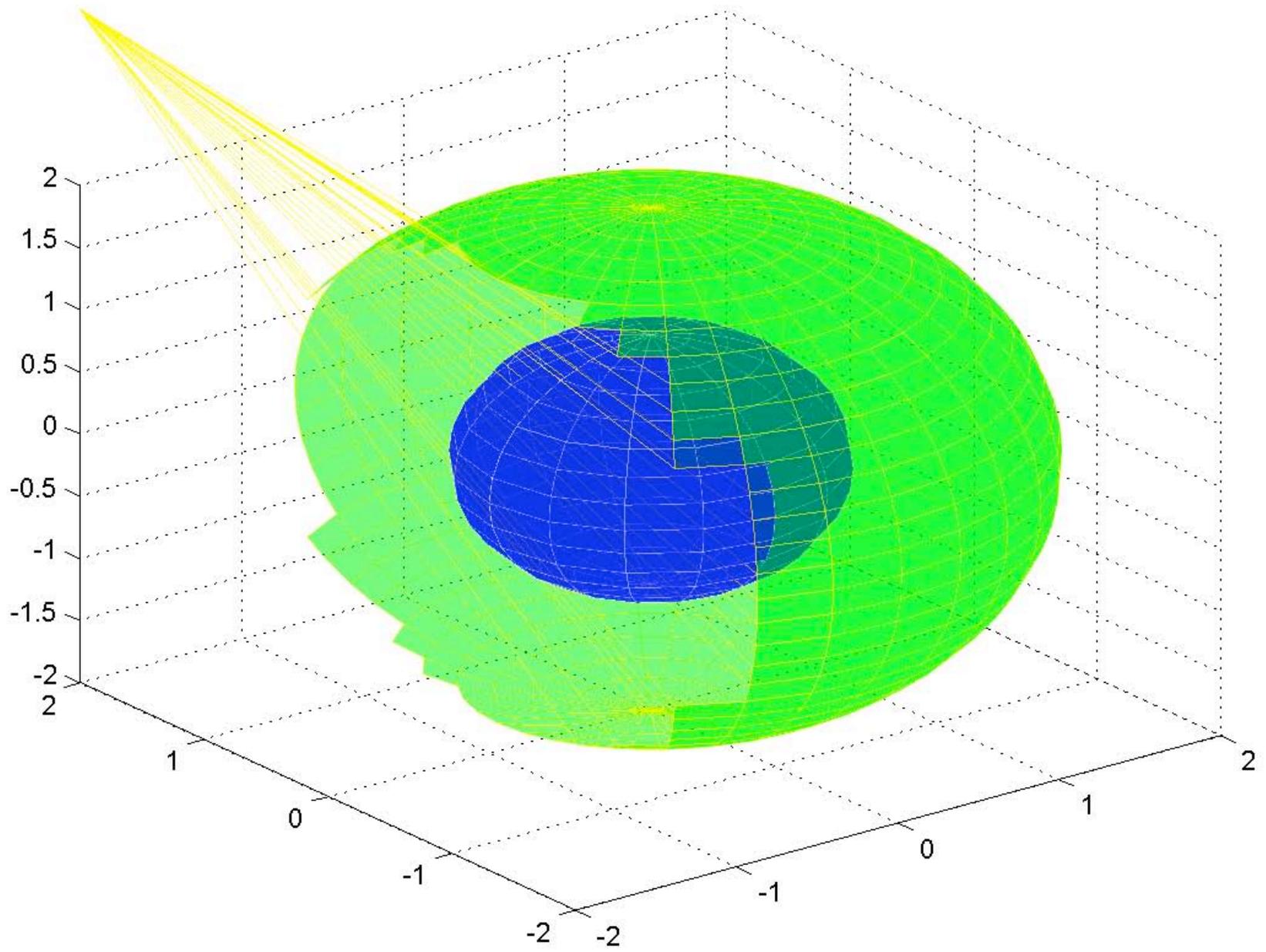
Photon Apportionment

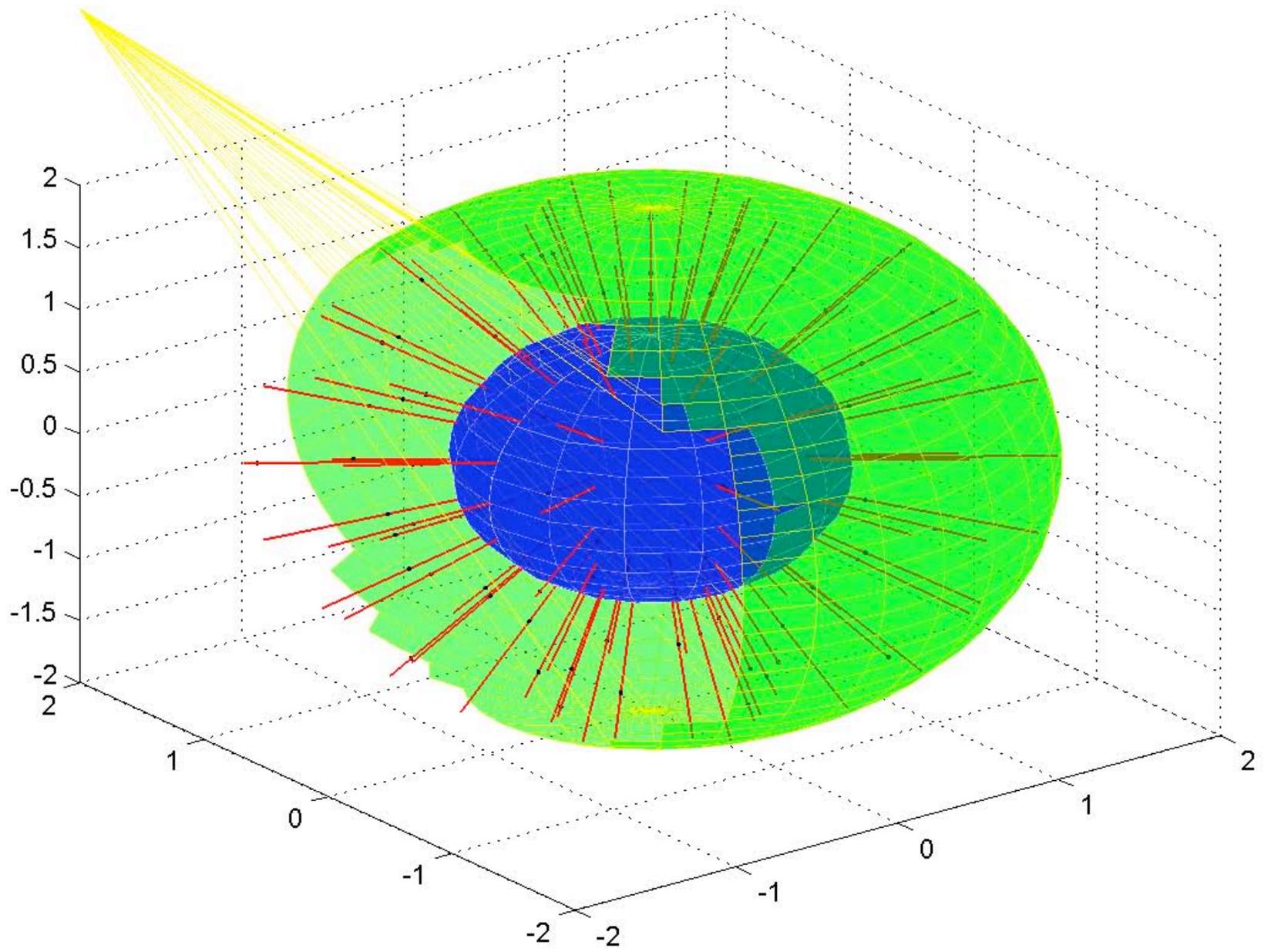


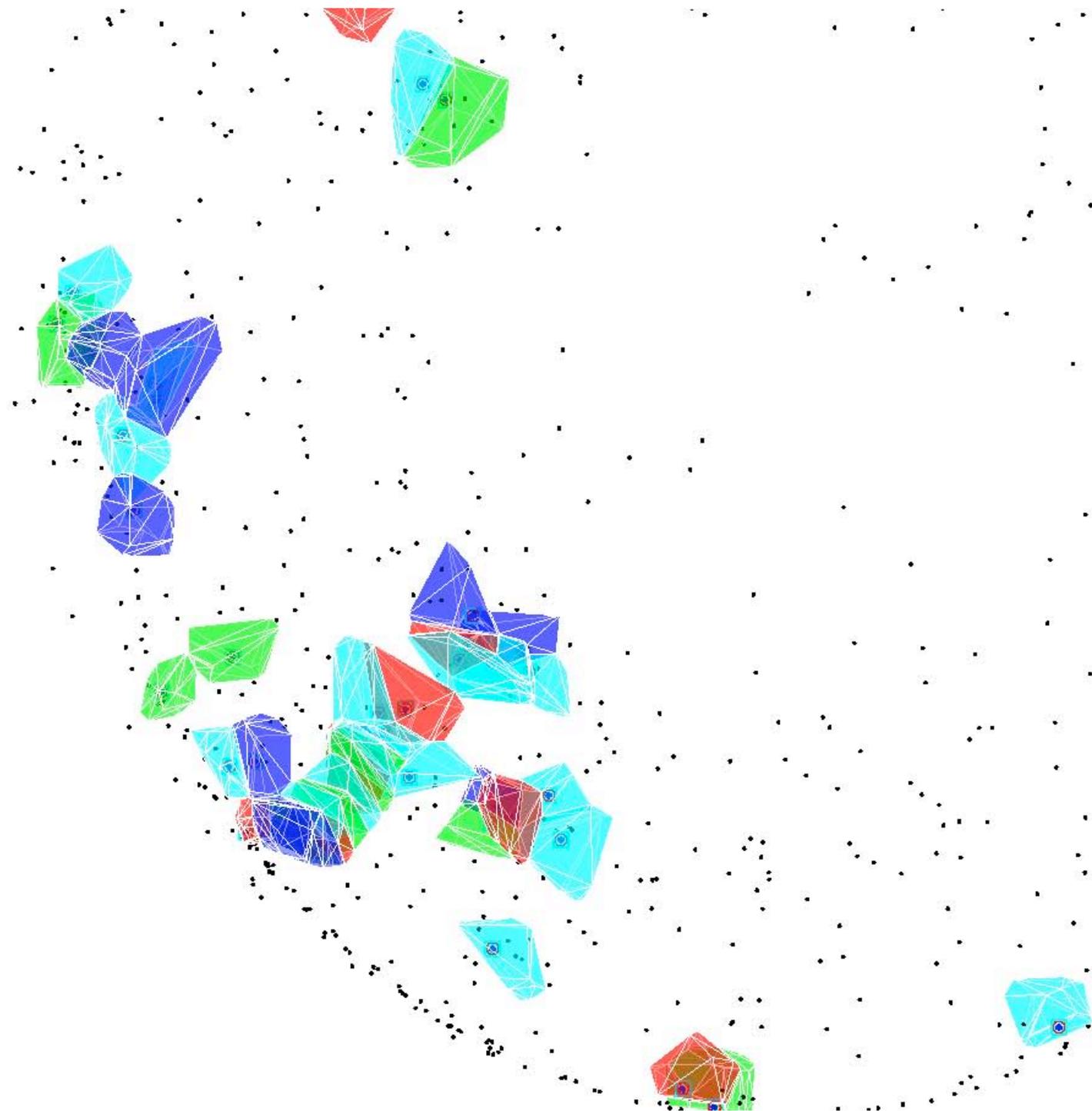
Differential: $\log(\text{PSF Contribution} / \text{Voronoi Area})$

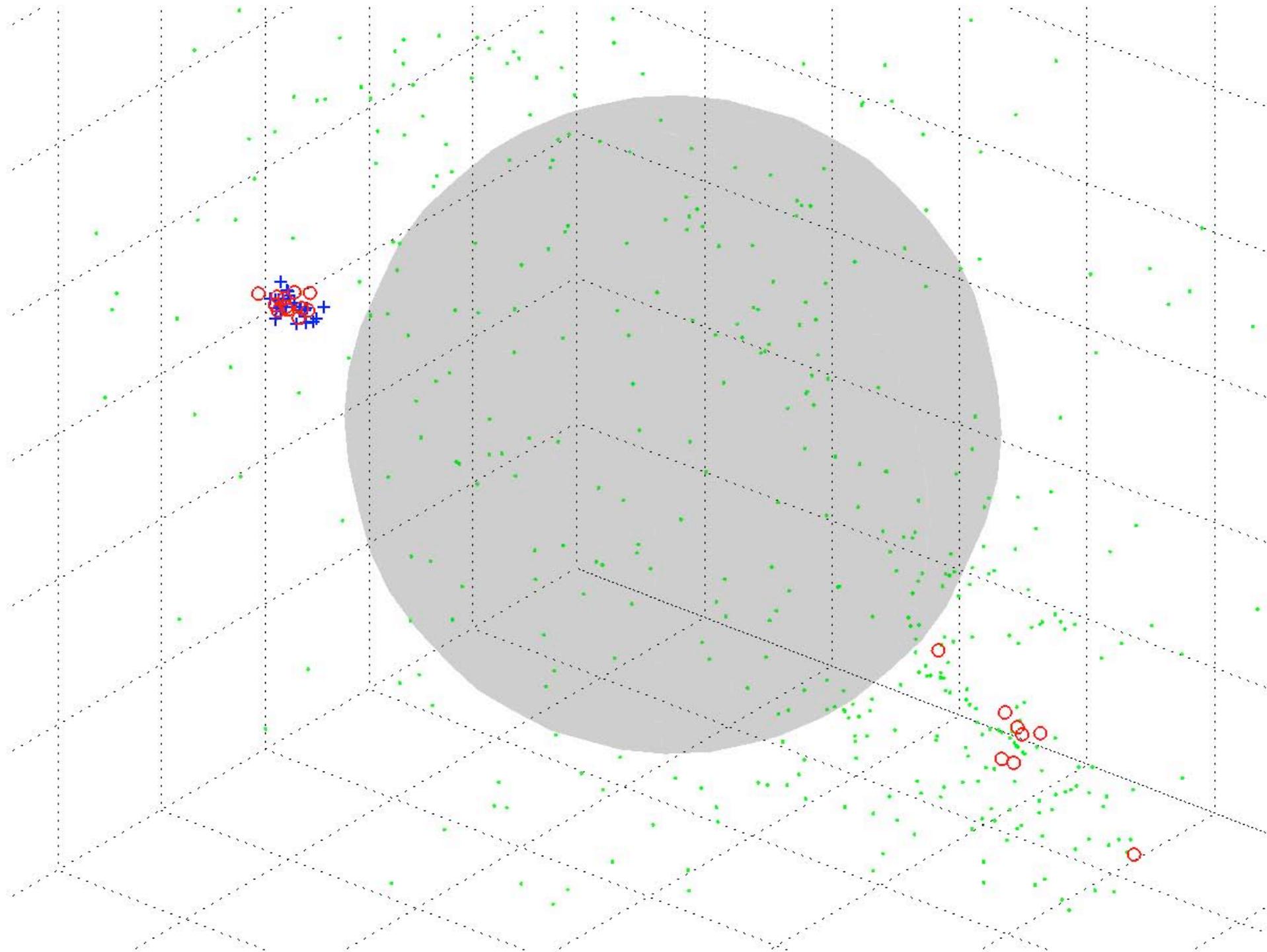














Features of the Algorithm

- No Bins (space, time, energy)
 - No Smoothing (space, time, energy)
- (No loss of information due to these approximations.
Result not dependent on bin sizes or locations.)
- Fast: $O(N)$
 - Incremental: $O(\Delta N)$ – (work by Giuseppe Romeo)
 - Suitable for quick look/automated science
 - No Coordinate Singularities on the Sphere
 - Flexible criterion for detection ...



Detection Criteria

Can evaluate the following at each photon:

- Local Density ($1 / \text{Voronoi volume}$)
- Clustering (connections to adjacent Voronoi cells)
- Difference in Spectrum (transient vs. background)
- Time difference (Voronoi volume vs. average of previous cell volumes nearby on the sky)
- Any other logically expressible criterion

... and incorporate them in the transient detection criterion.